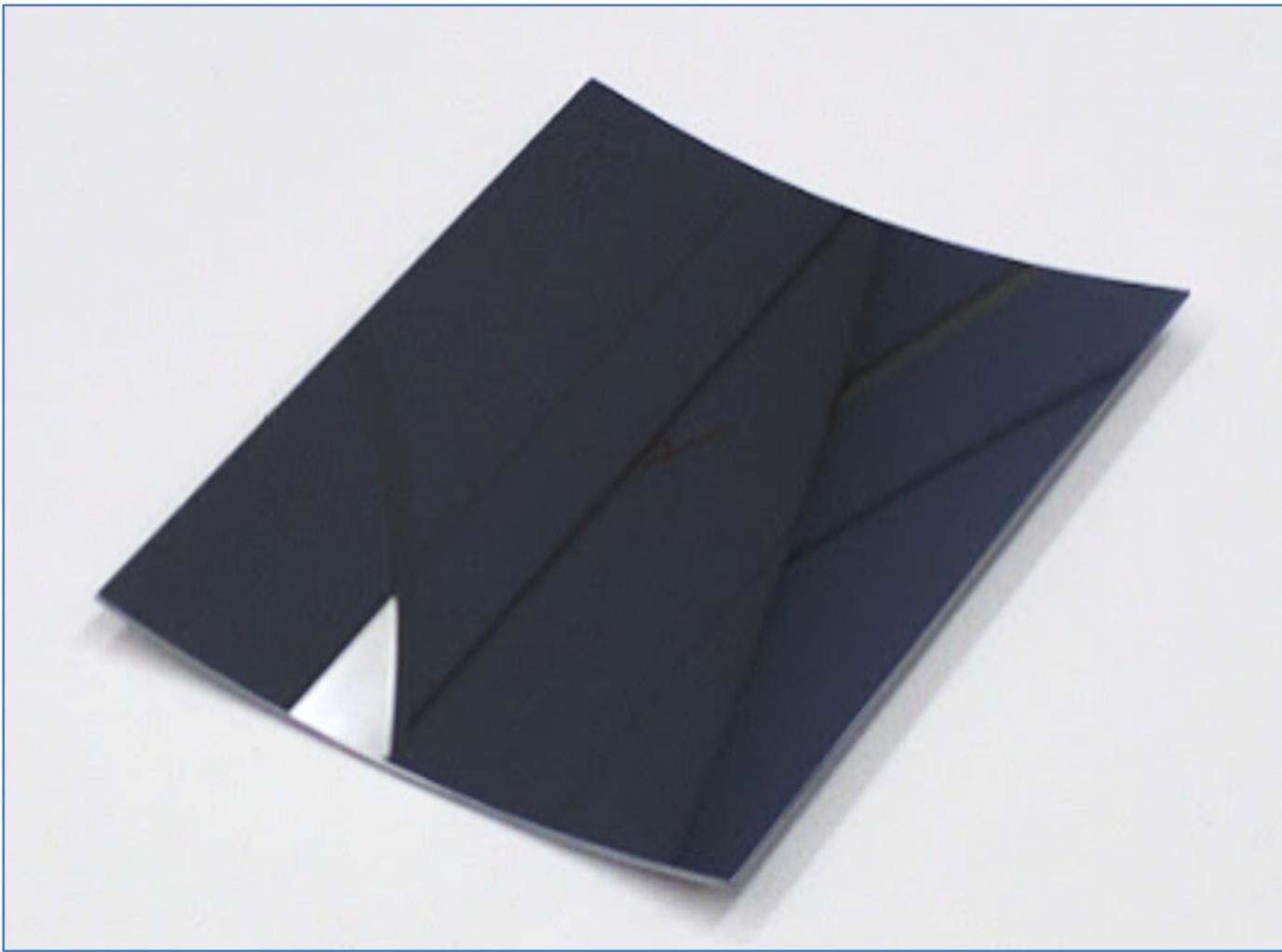




**2017 Hardware Images**

**PCOS and COR Strategic Technology Portfolio**

For more information about these technologies visit our Technology Database (<http://www.astrostrategictech.us>)

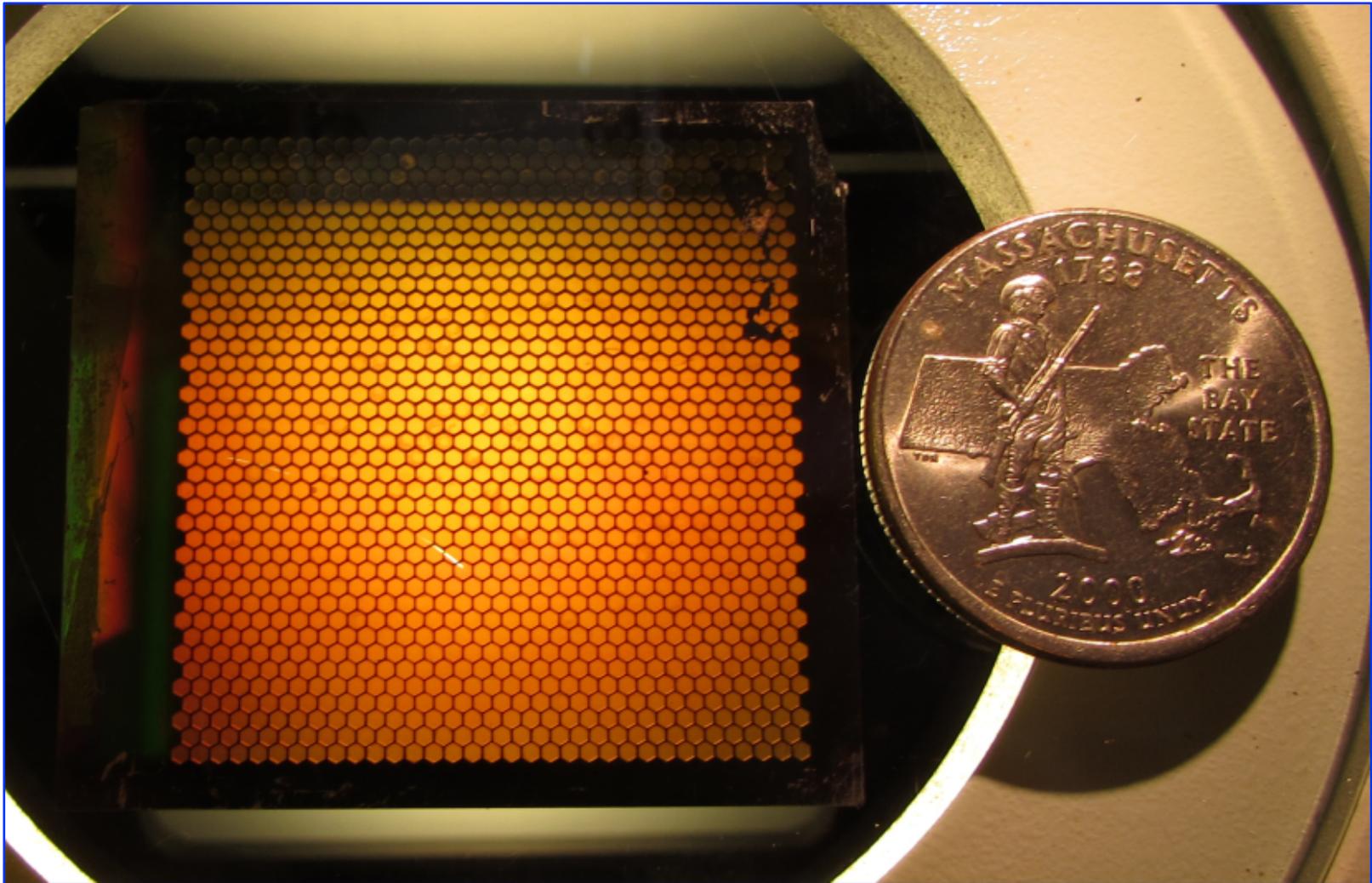


Thin single-crystal silicon X-ray mirror segment (after trimming)

**Significance:** World-class thin grazing-angle X-ray mirror technology; baselined for Lynx X-ray flagship mission concept

**Project Title:** Next Generation X-ray Optics: High Resolution, Light Weight, and Low Cost

**PI:** Zhang, William (GSFC)

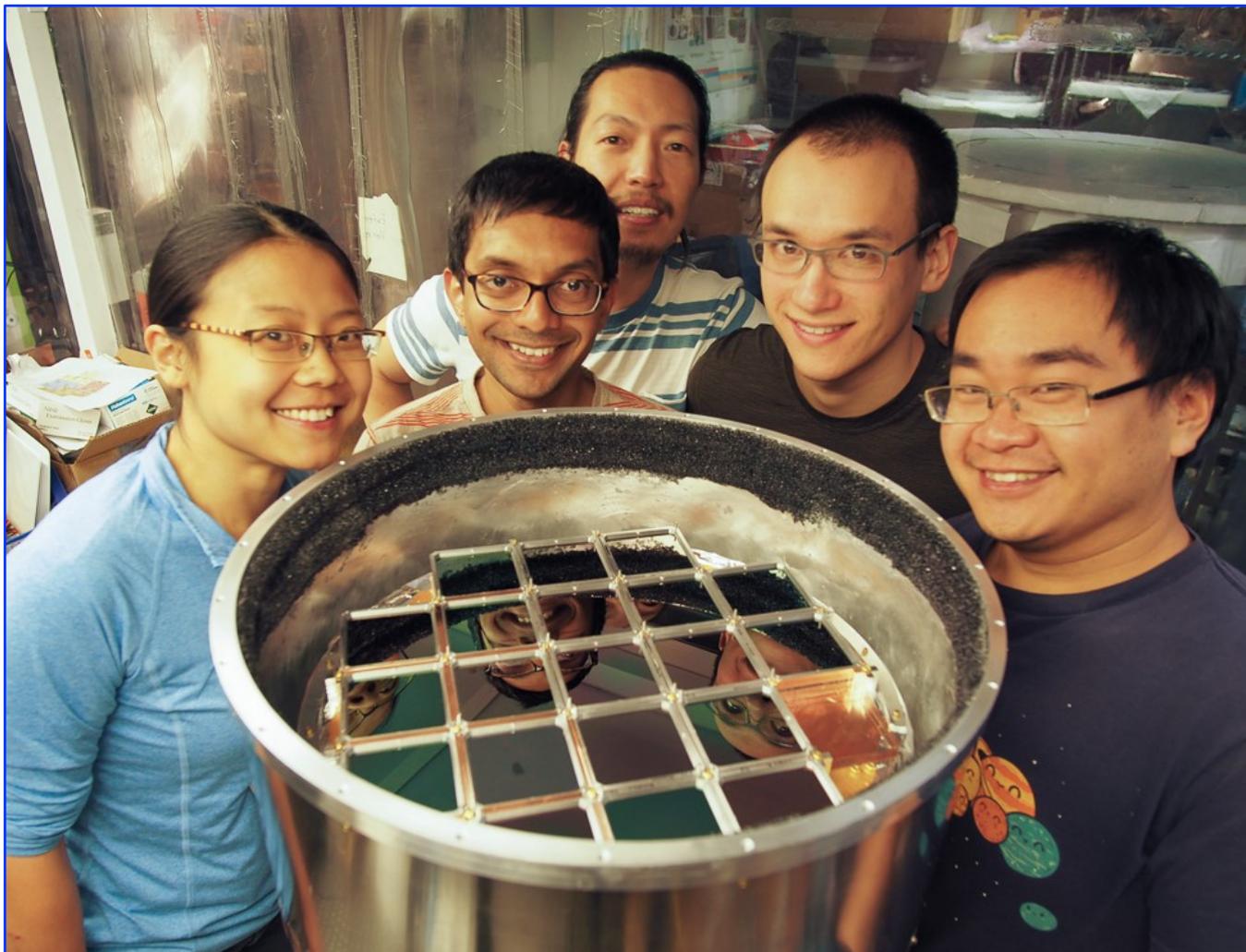


X-ray Critical-Angle Transmission (CAT) grating with quarter dollar coin for scale

**Significance:** Highest-resolution X-ray grating technology; baselined for Lynx X-ray flagship mission concept

**Project Title:** Development of a CAT Grating Spectrometer

**PI:** Mark Schattenburg (MIT Kavli Institute for Astrophysics and Space Research)

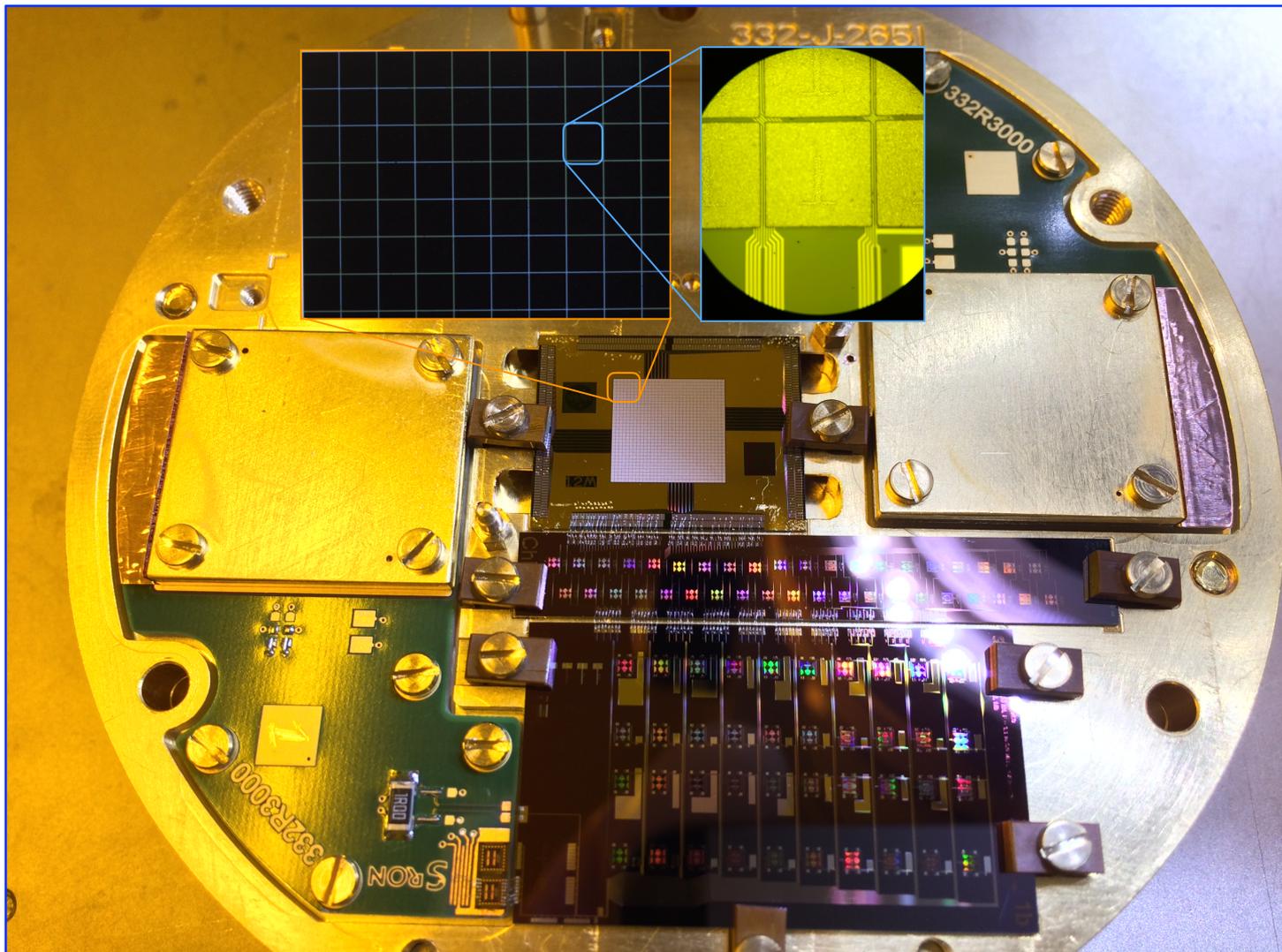


## Modular BICEP3 Focal Plane for Cosmic Microwave Background (CMB) polarimetry

**Significance:** Developing antenna designs providing sensitivity, stability, and minimized particle susceptibility for bands required by the Inflation Probe, enabling identification of Inflation instants after the Big Bang

**Project Title:** Planar Antenna-Coupled Superconducting Detectors for CMB Polarimetry

**PI:** James Bock (JPL/Caltech)

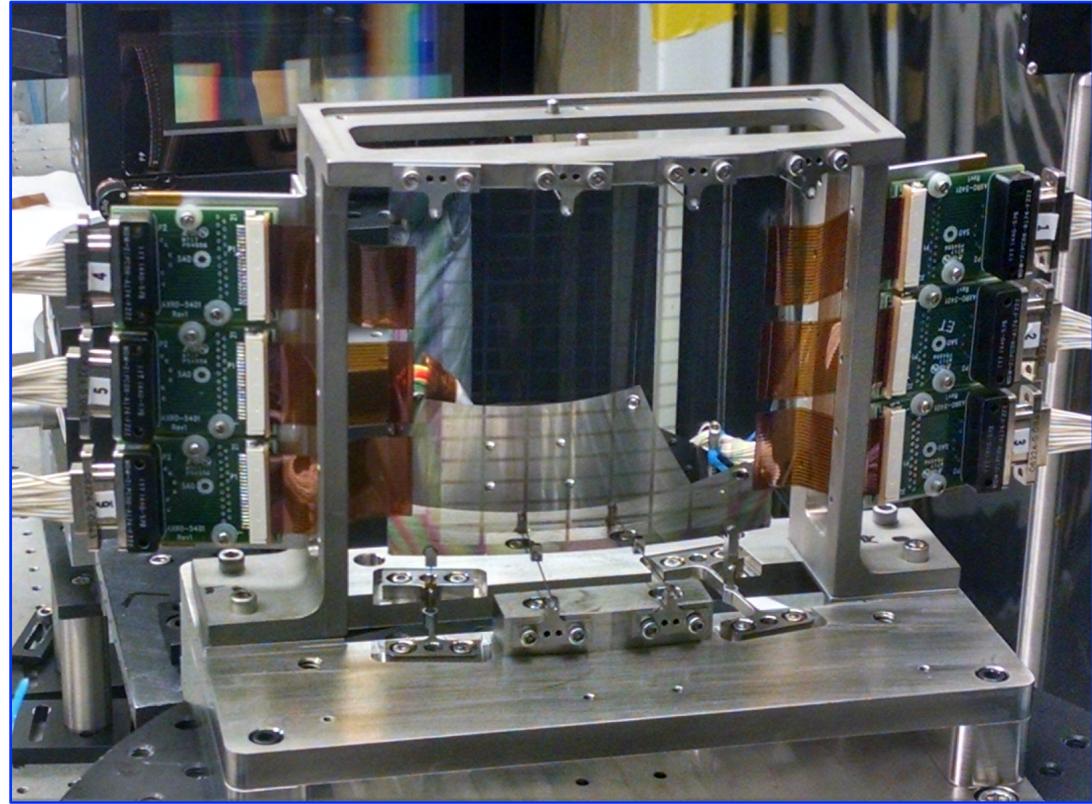
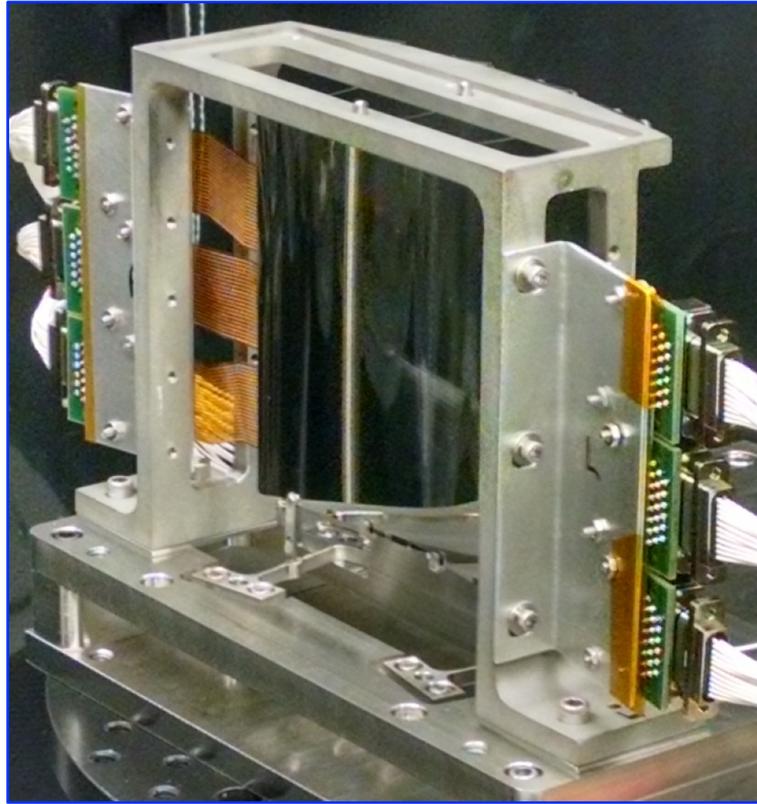


## Demonstration model array of Transition-Edge Sensors (TESs)

**Significance:** TES microcalorimeters offer energy resolution that may enable future missions such as the Lynx X-ray flagship mission concept

**Project Title:** Advanced X-ray Microcalorimeters: TES Microcalorimeters

**PI:** Caroline Kilbourne (GSFC)

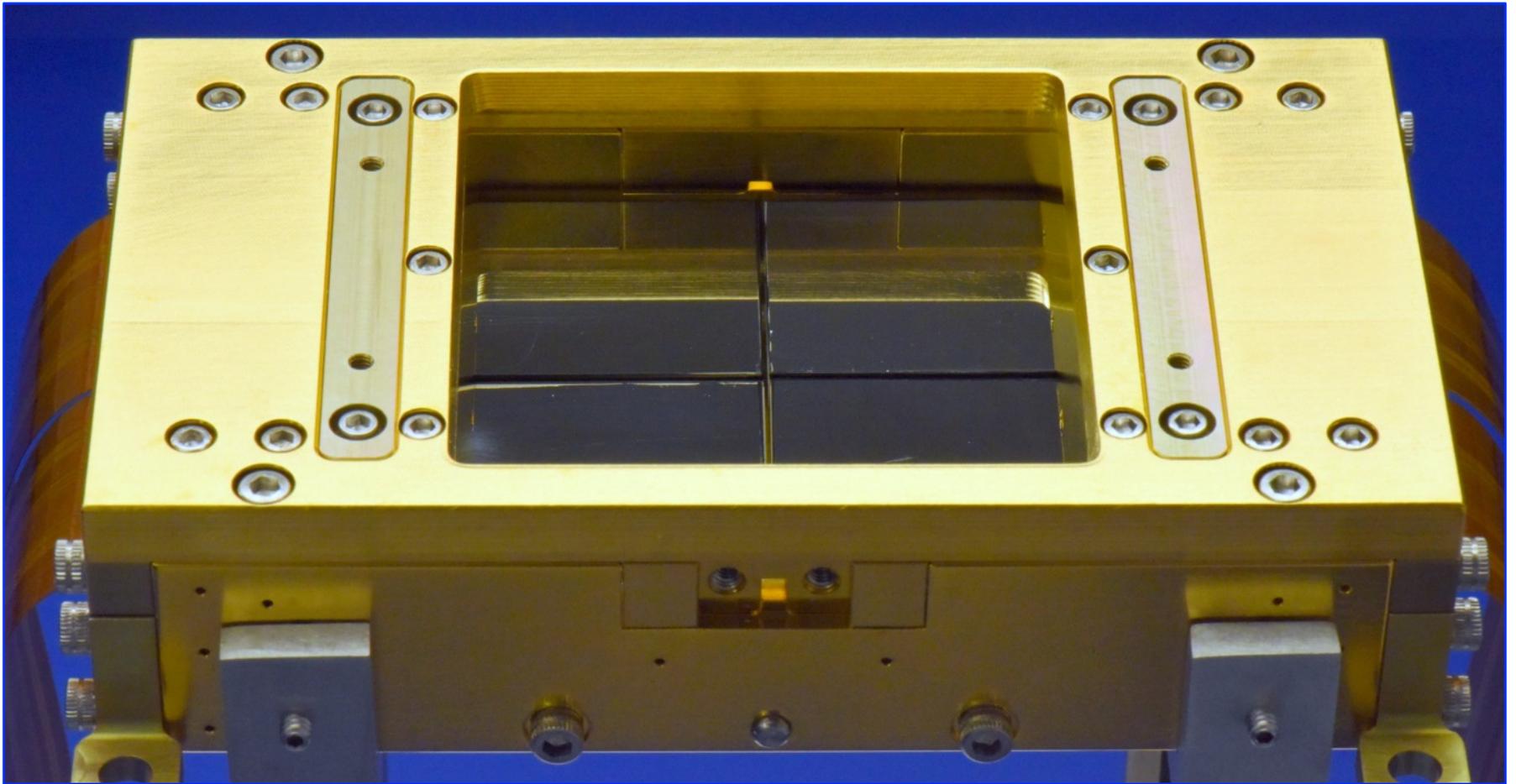


Single-shell mounting concept with mounted adjustable X-ray mirror and electrical connections

**Significance:** Adjustable X-ray optics are a backup technology for the Lynx X-ray large mission concept

**Project Title:** Adjustable X-Ray Optics

**PI:** Paul Reid (SAO)

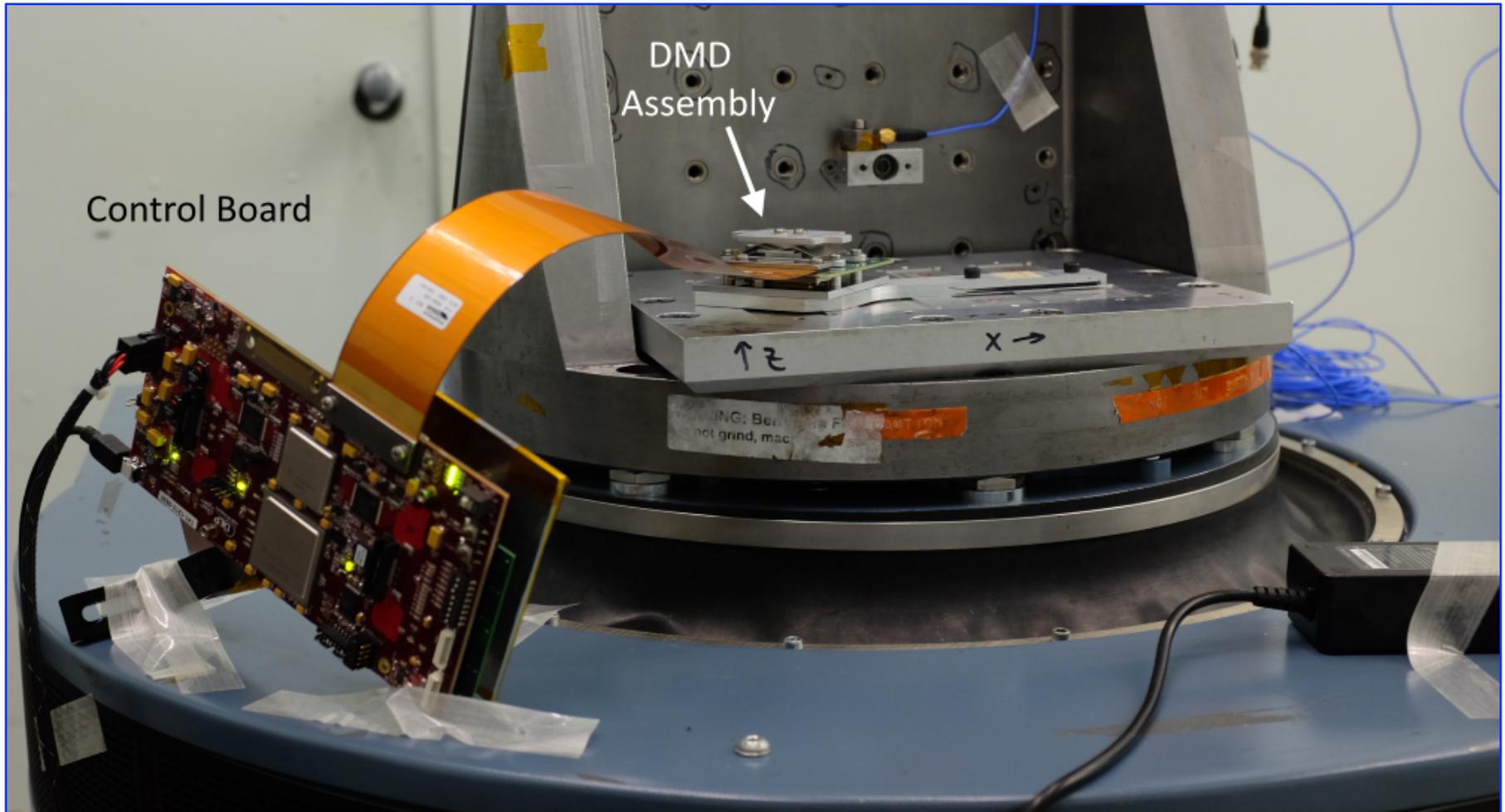


REXIS Detector Assembly Module (DAM) with directly deposited filter on the X-ray CCDs; REXIS is a student-built instrument deployed on the OSIRIS-REx mission to Asteroid Bennu

**Significance:** X-ray detectors operate far better when filters allow X-ray photons through and block longer wavelength light

**Project Title:** Directly-Deposited Blocking Filters for X-ray Imaging Detectors

**PI:** Mark Bautz (MIT Kavli Institute for Astrophysics and Space Research)



Digital Micro-mirror Device (DMD) undergoing vibration testing as part of flight qualification

**Significance:** Replacing windows of commercially available DMDs may enable far-UV multi-object spectrometry in future missions

**Project Title:** Development of DMDs for Far-UV Applications

**PI:** Zoran Ninkov (RIT)

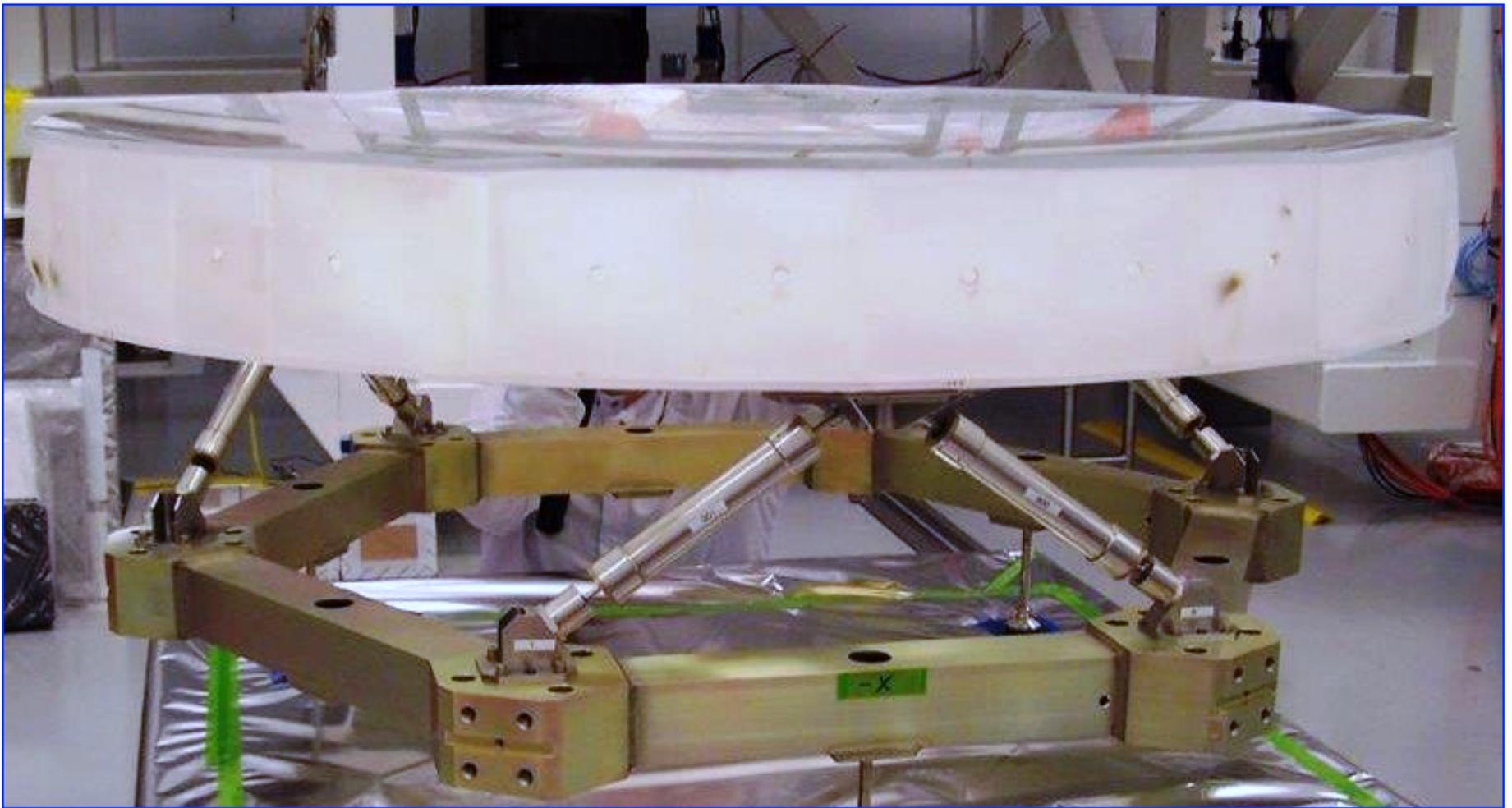


Selected parts for 4-to-0.05-K for Continuous Adiabatic Demagnetization Refrigerator (CADR) cooling system fabricated at GSFC

**Significance:** This advanced sub-Kelvin cooling technology has been baselined by Lynx, Origins, PICO, and GEP

**Project Title:** High-Efficiency Continuous Cooling for Cryogenic Instruments and sub-Kelvin Detectors

**PI:** James Tuttle (GSFC)

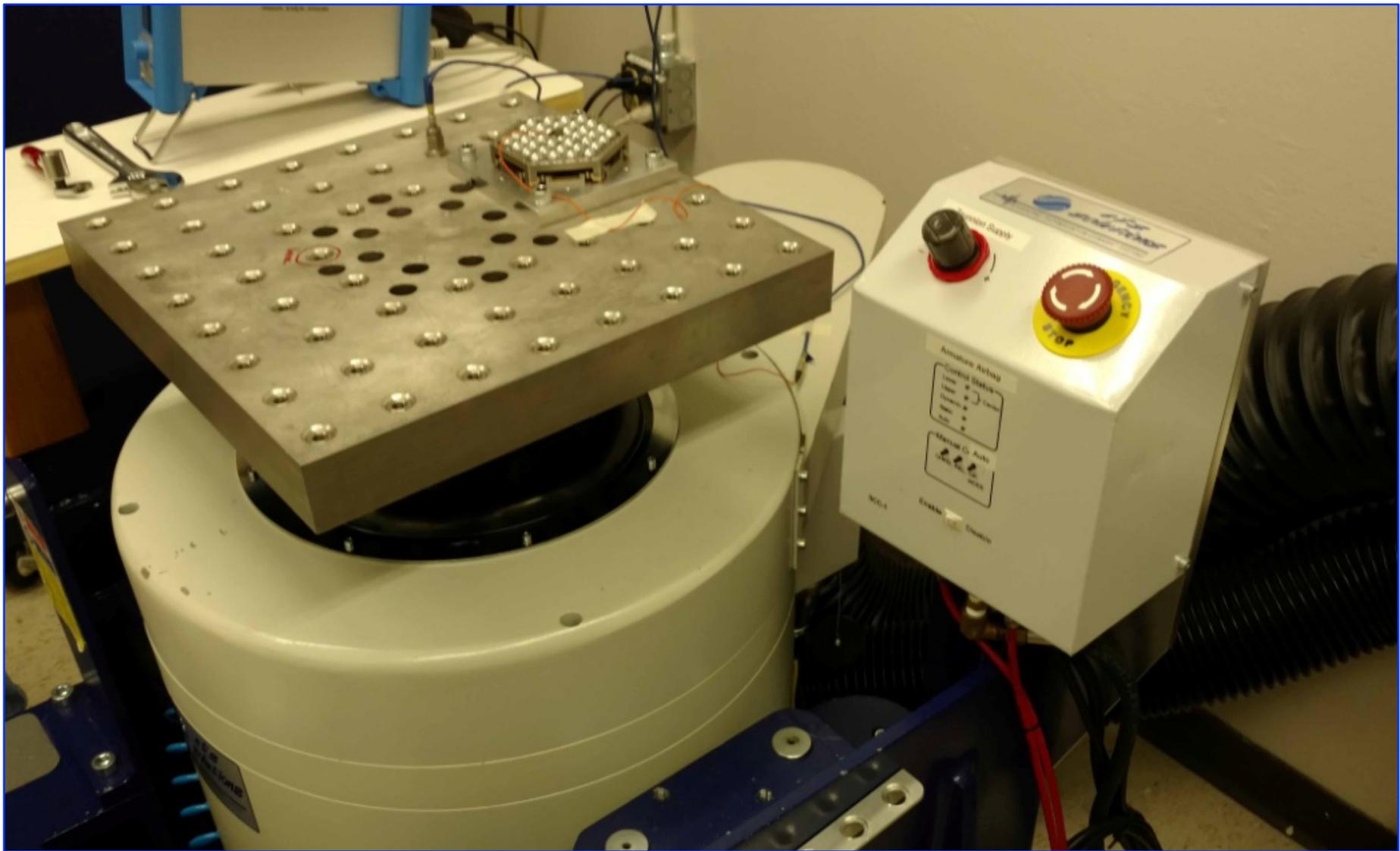


### 1.5-m Harris mirror with support structure

**Significance:** This technology may enable required ultra-stability ( $\sim 10$  pm) for HabEx and LUVOIR missions

**Project Title:** Predictive Thermal Control (PTC) Technology to enable Thermally Stable Telescopes

**PI:** H. Philip Stahl (MSFC)

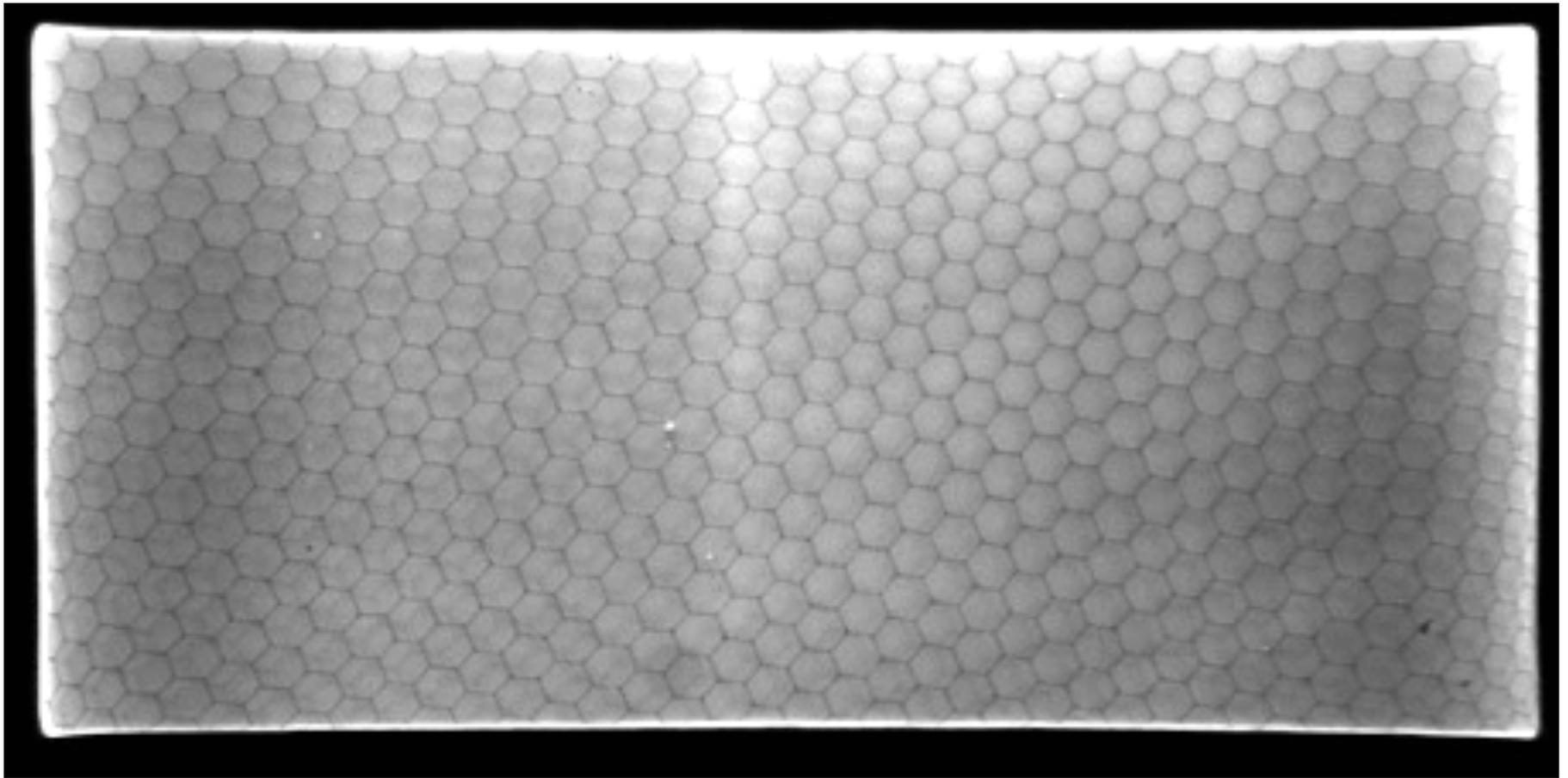


First resonance measurements of mock Cosmic Microwave Background (CMB) detector array

**Significance:** This and related technologies may enable future CMB missions, e.g. LiteBIRD

**Project Title:** Technology Development for LiteBIRD and other CMB Missions

**PI:** Adrian T. Lee (UC Berkeley)

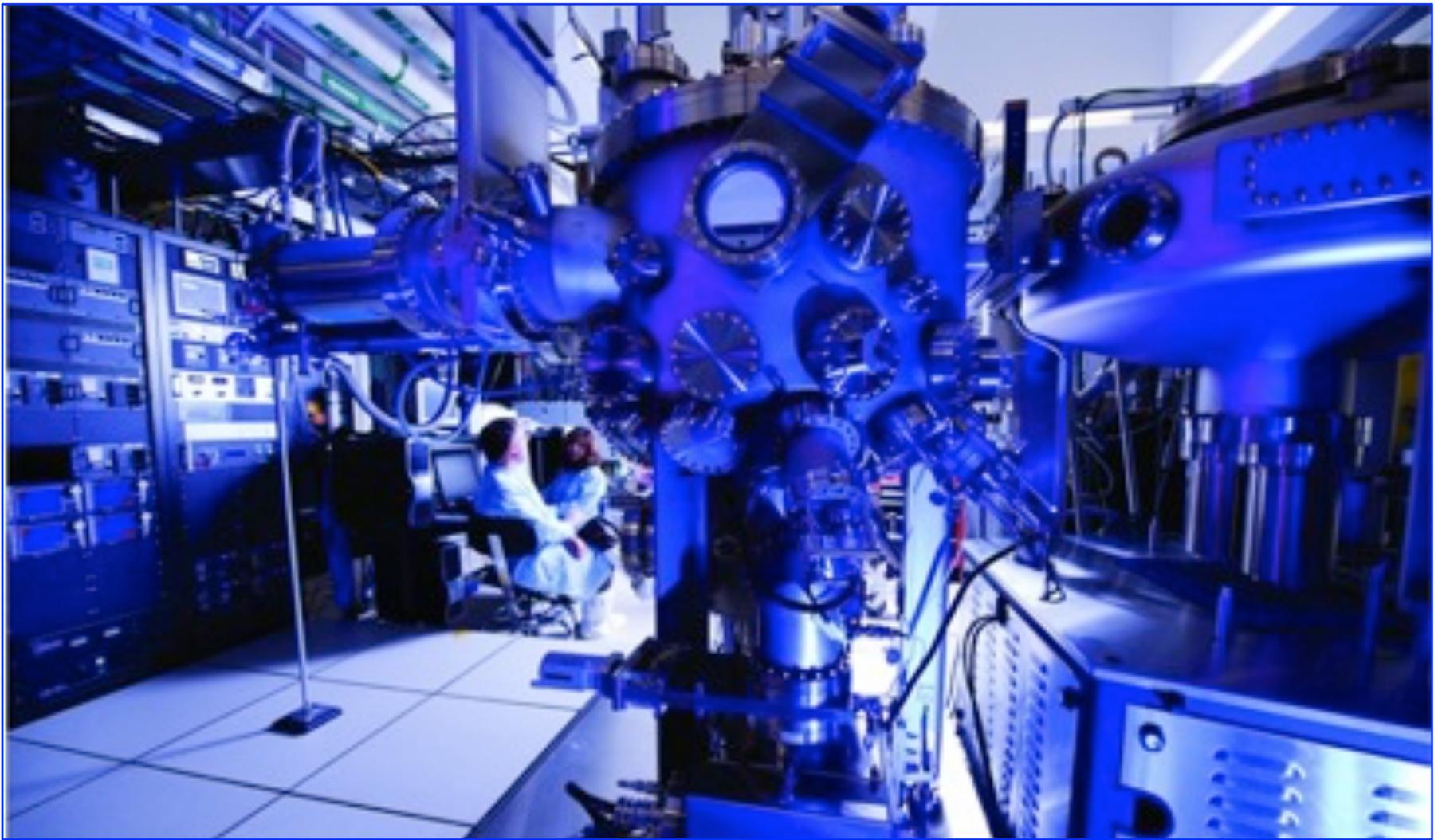


### [Atomic Layer Deposition \(ALD\) Multi-Channel Plate \(MCP\) image detail](#)

**Significance:** This detector technology is baselined by HabEx, LUVOIR, and CETUS for UV/Visible light detection

**Project Title:** High-Performance Sealed-Tube Cross-Strip (XS) Photon-Counting Sensors for UV-Vis Astrophysics Instruments

**PI:** Oswald Siegmund (UC Berkeley)

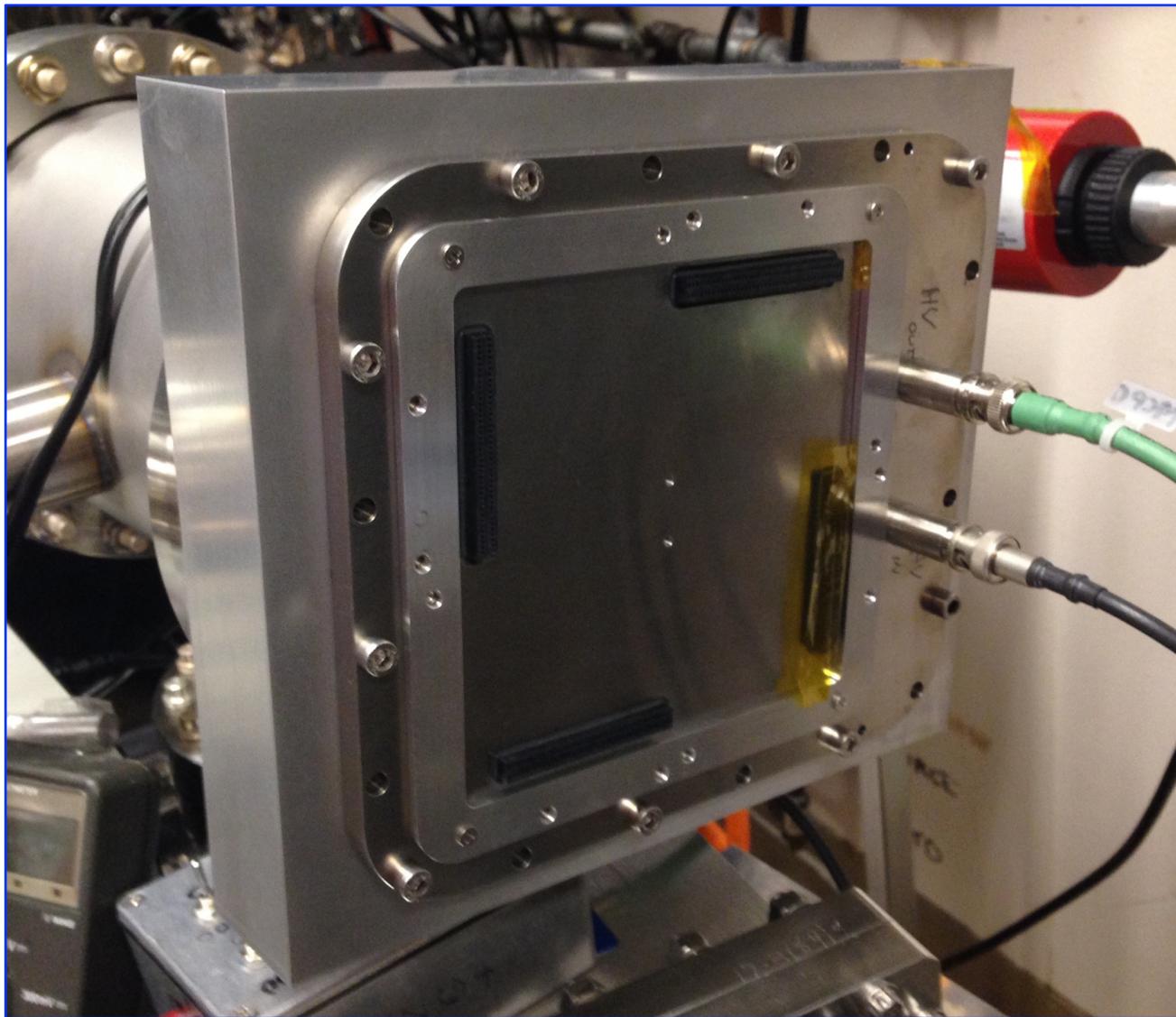


**Molecular Beam Epitaxy (MBE) system at JPL used for coating advanced detectors**

**Significance:** Advanced detectors developed by this project are baselined by SHIELDS, HabEx, LUVOIR, and ground facilities are fabricated using Atomic Layer Deposition (ALD) coatings

**Project Title:** Advanced FUV/UV/Visible Photon-Counting and Ultralow-Noise Detectors

**PI:** Shouleh Nikzad (JPL/Caltech)



100×100 mm<sup>2</sup> Multi-Channel Plate (MCP) detector mounted on vacuum test system

**Significance:** Large-format low-noise detectors may enable future far-UV missions

**Project Title:** Development of 100×100 mm<sup>2</sup> photon-counting UV detectors

**PI:** John Vallergera (UC Berkeley)

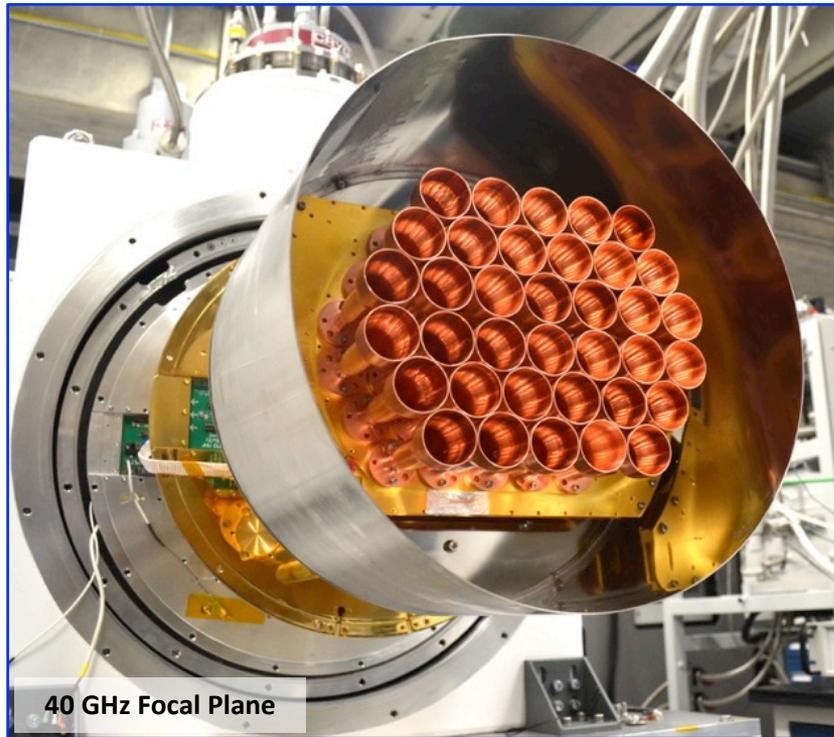
Stratospheric Terahertz Observatory (STO2)  
Balloon-borne observatory being prepared for launch in Antarctica; an early working device was provided by this team to fly on STO2 in 2016

**Significance:** This technology provides 4.7-THz Local Oscillators (LOs), enabling far-IR/sub-mm missions such as the balloon-borne Galactic/Extragalactic ULDB Spectroscopic Terahertz Observatory (GUSTO)

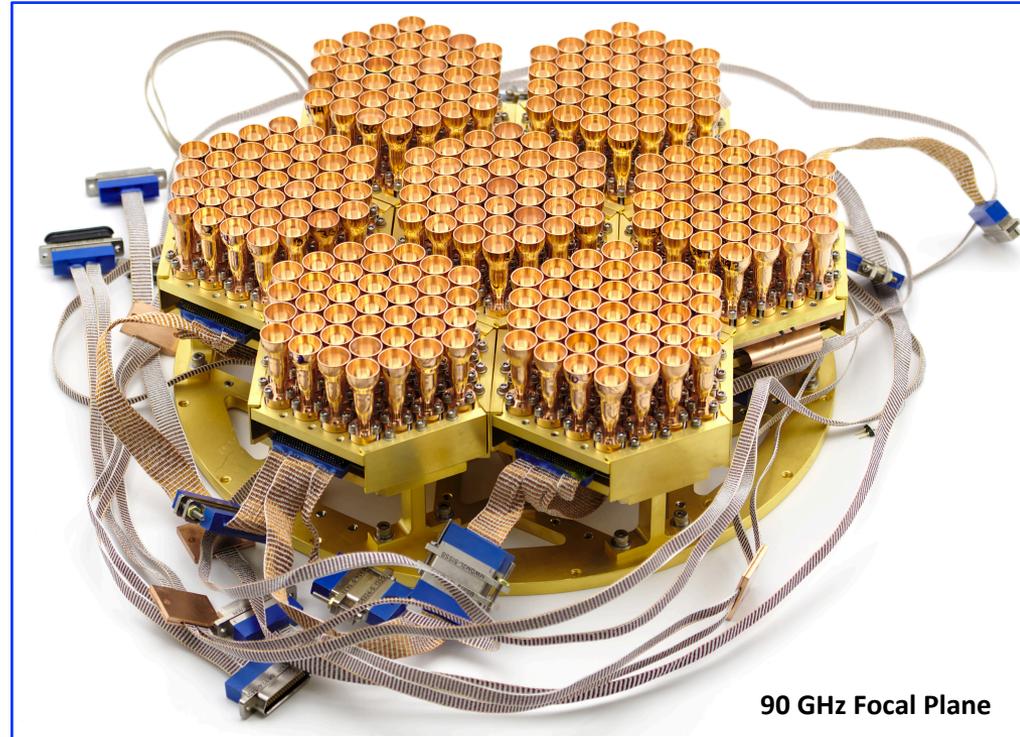
**Project Title:** Raising the Technology Readiness of 4.7-THz local oscillators

**PI:** Qing Hu (MIT)





40 GHz Focal Plane



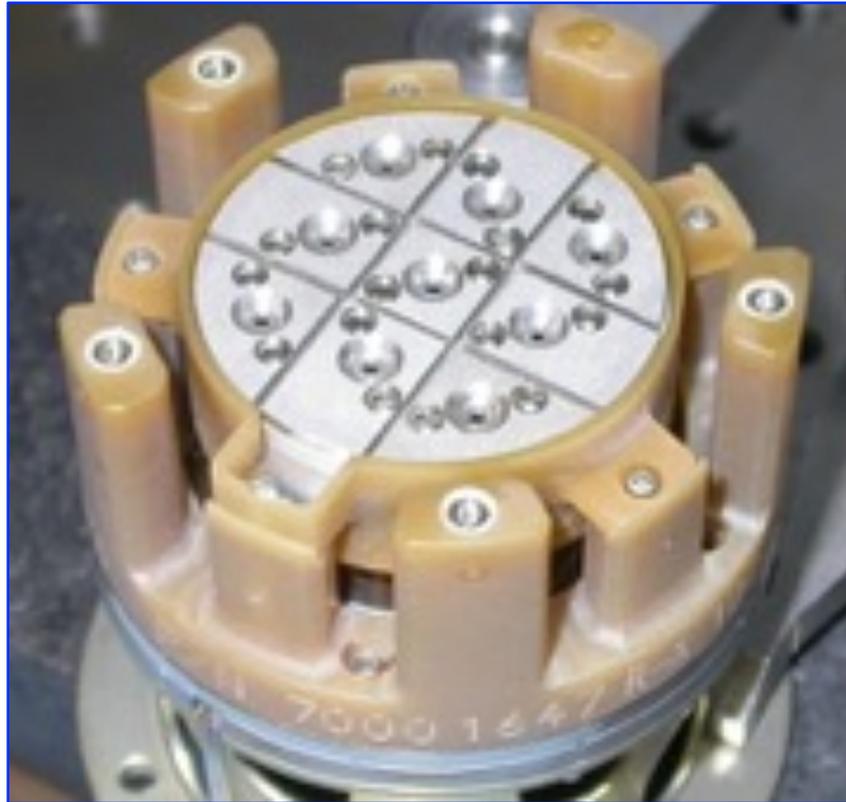
90 GHz Focal Plane

## 40-GHz and 90-GHz feedhorn-coupled focal plane arrays for performing Cosmic Microwave Background (CMB) measurements

**Significance:** CMB measurements may enable identification of the “Inflation” cosmologists believe may have occurred instants after the Big Bang; multi-wavelength measurements can help remove foreground noise

**Project Title:** High Efficiency Feedhorn-Coupled TES-based Detectors for CMB Polarization

**PI:** Edward Wollack (GSFC)

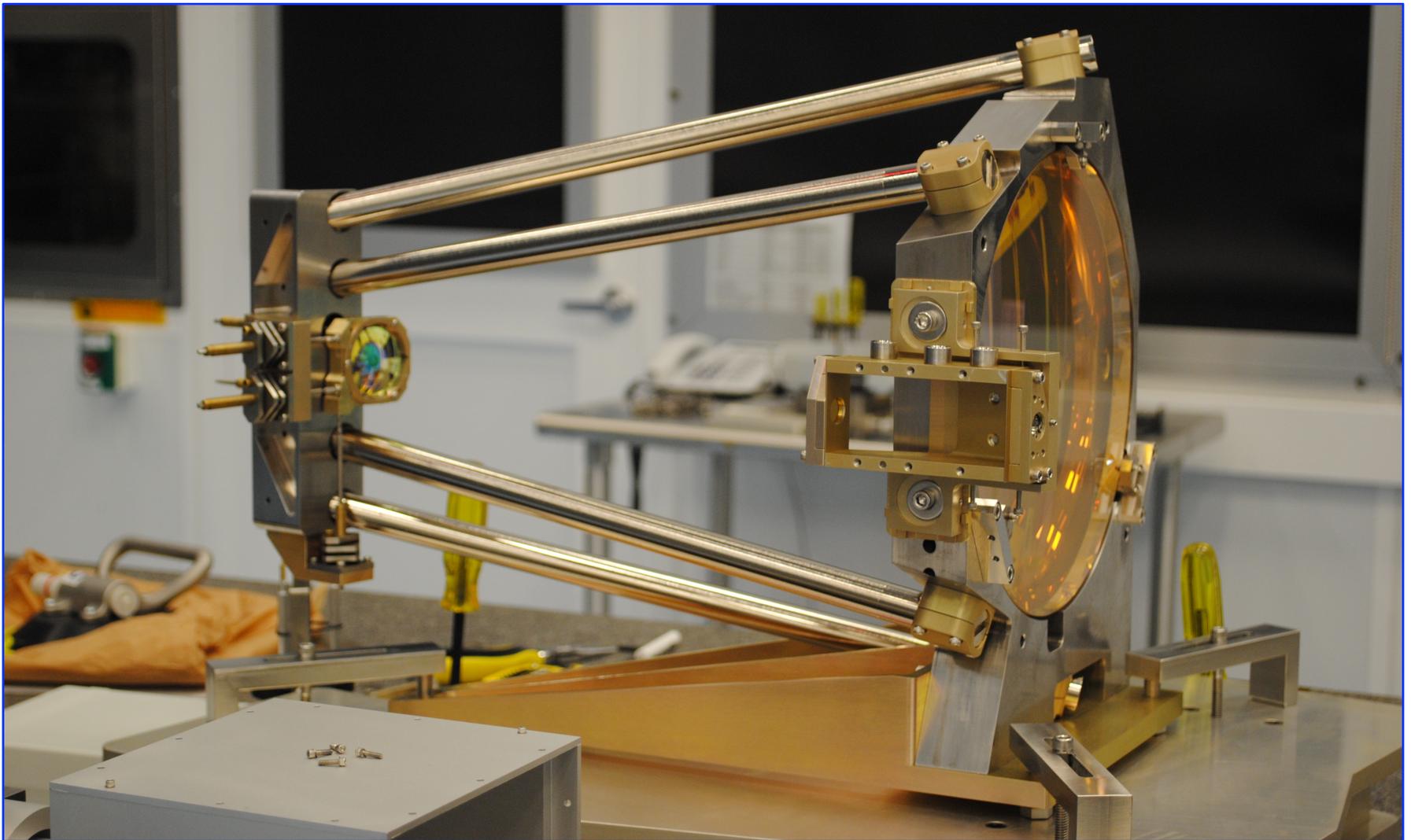


[Busek components developed for colloid microthrusters for precision spacecraft control](#)

**Significance:** Microthrusters are critical for controlling the spacecraft of future missions such as the Laser Interferometer Space Antenna (LISA) gravitational-wave observatory

**Project Title:** LISA Colloid Microthruster Technology

**PI:** John Ziemer (JPL)

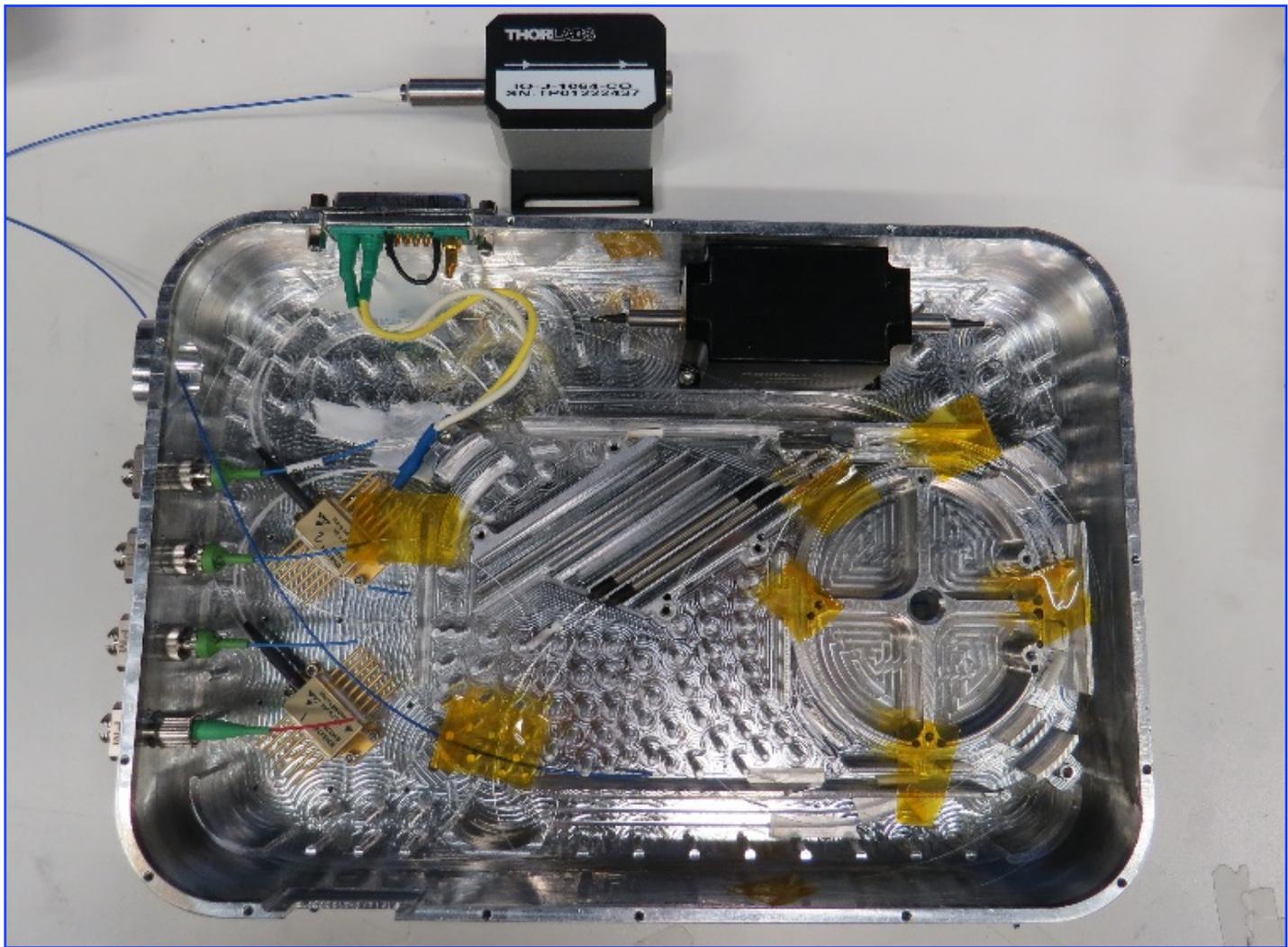


Prototype Laser Interferometer Space Antenna (LISA) telescope designed for room-temp testing

**Significance:** The LISA gravitational-wave observatory crucially depends on collecting laser light from a remote spacecraft, millions of km away

**Project Title:** Telescope Development for the LISA Mission

**PI:** Jeffrey Livas (GSFC)



Prototype laser power amplifier for lasers enabling the Laser Interferometer Space Antenna (LISA) gravitational-wave observatory

**Significance:** LISA crucially depends on lasers to allow interferometric measurement of the multi-million-km distance between the three spacecraft

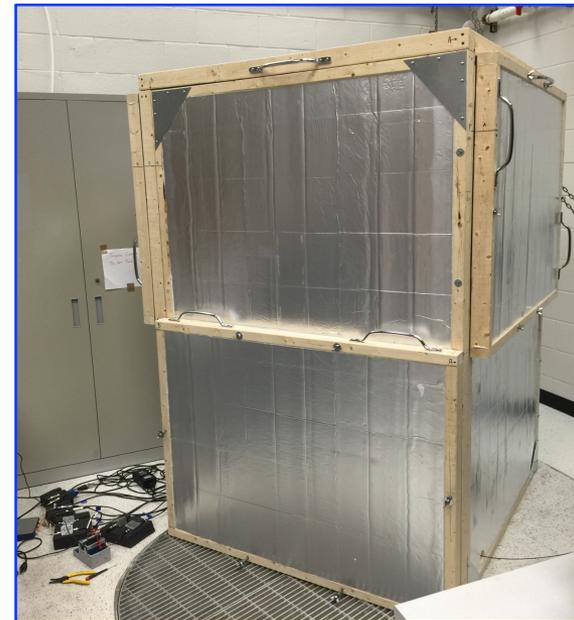
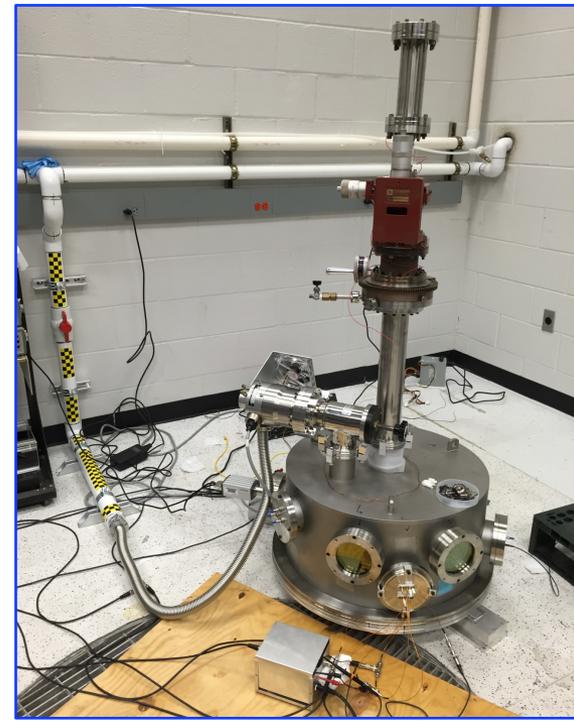
**Project Title:** Laser Technology Development Project for the LISA Mission

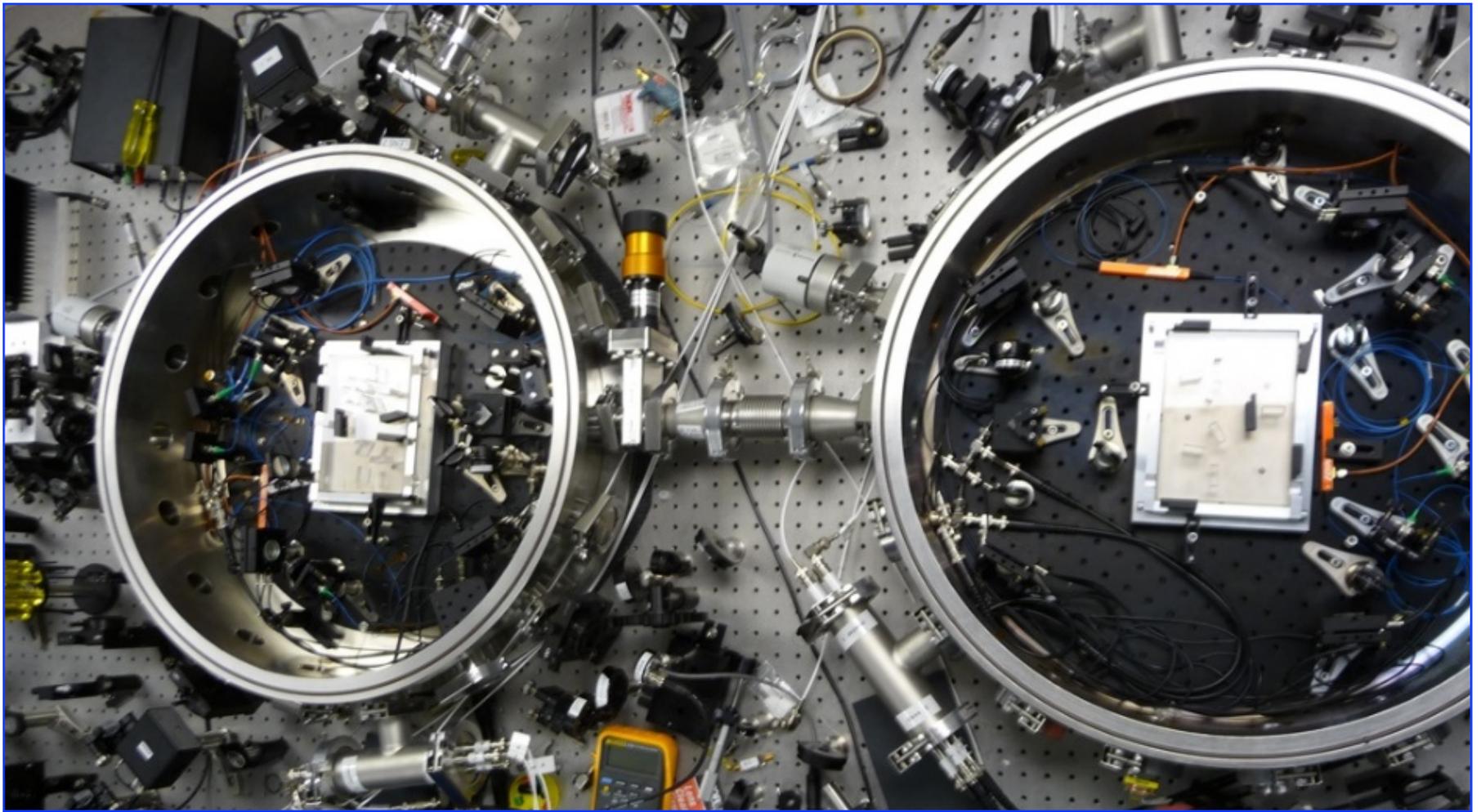
**PI:** Anthony Yu (GSFC)

Torsion pendulum test-bed for gravitational reference sensors used for testing charge management system for the Laser Interferometer Space Antenna (LISA) gravitational-wave observatory

**Significance:** LISA crucially depends on charge management to prevent electrostatic noise interfering with interferometric measurement of the multi-million-km distance between the three spacecraft

**Project Title:** UV LED-based Charge Management System for LISA  
**PI:** John Conklin (University of Florida)



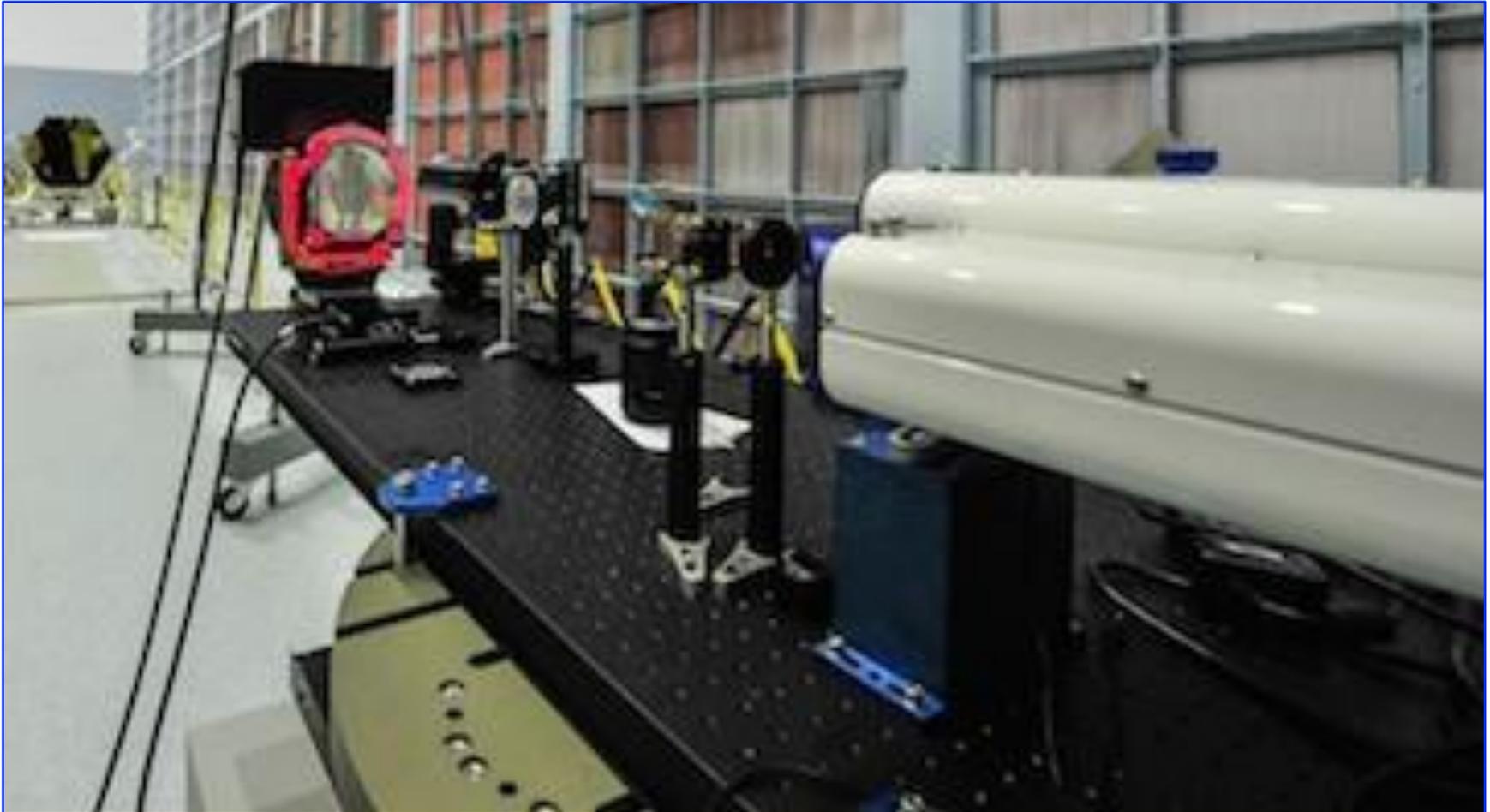


Interferometer testbed built to demonstrate phasemeter and measurement system performance developed for the Laser Interferometer Space Antenna (LISA) gravitational-wave observatory

**Significance:** LISA needs a phasemeter system to allow interferometric measurement of the multi-million-km distance between the three spacecraft

**Project Title:** LISA Phasemeter Project

**PI:** William Klipstein (JPL)



Lab test setup for picometer measurement of mirror dynamics

**Significance:** Ultra-stability and -precision ( $\sim 10$  pm) may enable the HabEx and LUVOIR missions

**Project Title:** Ultra-Stable Structures

**PI:** Babak Saif (GSFC)