2023 Hardware Images PhysCOS and COR Strategic Technology Portfolio

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An X-ray mirror module on a vibration table

Significance: World-class thin grazing-angle X-ray mirror technology that would enable the next X-ray Great Observatory

Project Title: Next Generation X-ray Optics: High Resolution, Light Weight, and Low Cost **PI:** Zhang, William (GSFC)



Critical-Angle Transmission (CAT) gratings being tested at the PANTER X-ray facility Significance: Highest-resolution X-ray transmission grating technology that could fly on the next X-ray Great Observatory Project Title: Technology Maturation for a High-Sensitivity and High-Resolving-Power X-ray

Spectrometer

PI: Mark Schattenburg (MIT Kavli Institute for Astrophysics and Space Research) **Co-I:** Ralf Heilmann



Four Critical-Angle Transmission (CAT) grating facets aligned and mounted for testing
 Significance: Demonstrates manufacturability of highest-resolution X-ray transmission
 grating technology that could fly on the next X-ray Great Observatory
 Project Title: Readying X-ray Gratings and Optics for Space Applications: Manufacturability
 and Alignment
 PI: Randall Smith (SAO)



16.7-MPixel CMOS image sensor
Significance: Low-noise detectors are crucial for future missions
Project Title: A Single-Photon-Sensing and Photon-Number-Resolving Detector for NASA Missions
PI: Don Figer (RIT)



Detailed images of a Thermal Kinetic Inductance Detector (TKID)

Significance: Cosmic Microwave Background (CMB) polarimetry is crucial for understanding early universe physics. This project aims to ready NASA for the Inflation Probe identified by the 2020 Decadal Survey.

Project Title: Superconducting Detectors for CMB Polarimetry in PICO

PI: Roger O'Brient (JPL/Caltech)



X-ray CCD with Multi-Channel Readout Chip (MCRC) ready for testing Significance: Advanced X-ray detectors may enable the next X-ray Probe or Great Observatory Project Title: Extremely Low-noise, High Frame-rate X-ray Image Sensors for Strategic Astrophysics Missions PI: Mark Bautz (MIT Kavli Institute for Astrophysics and Space Research)



LEM 6-channel prototype readout for Transition-Edge-Sensor (TES) arrays

Significance: High-multiplexing-factor readouts may enable missions such as the next X-ray Great Observatory

Project Title: Microwave SuperconductingQUantum Interference Device (SQUID) Multiplexingfor Future X-ray Astrophysics MissionsPI: Douglas Bennett (NIST)



An ADR salt pill suspended in the bore of a 3-Tesla superconducting magnet via a lowthermal-conductivity suspension

Significance: This advanced sub-Kelvin cooling technology may enable multiple future strategic missions

Project Title: Development of ultra-low-temperature Continuous Adiabatic Demagnetization
 Refrigerator (CADR) with a Continuous Intermediate Stage for Heat Intercept
 PI: Mark Kimball (GSFC)

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Fanout board with superconducting flex lines, accommodating a 32 x 16 detector array and SQUID multiplexer for array readout. The board has a flexible region between two rigid ones. The rigid regions have indium bumps, to which a 2d-SQUID mux chip and a detector chip will be flip chip bonded. Having no vias simplifies board fabrication and ensures good reliability. **Significance:** Advanced far-IR detectors may enable the next Far-IR Great Observatory **Project Title:** Demonstrating Large, Low-Noise, Transition-Edge-Sensor Arrays for Future Far-IR Space Missions

PI: Johannes Staguhn (JHU & GSFC)



Ultra-High-Vacuum research chamber capable of reactive thin-film physical vapor deposition (rPVD) and passivation process **Significance:** High far-UV reflectance is hindered by oxidation of aluminum mirrors; preventing it may enable future far-UV missions

Project Title: Advanced Aluminum Mirrors with Passivated LiF for Environmentally Stable 1-m-Class UV Space Telescopes **PI:** Manuel Quijada (GSFC)



Using Computer-Generated Hologram (CGH) to absolutely characterize gravity-sag of mirrors

Significance: Ultra-high resolution general astrophysics observations require that the next IR/Optical/UV Great Observatory be diffraction- limited at 500 nm or better, this requires a primary mirror whose gravity-sag is known absolutely to a few nanometers rms.

Project Title: UV/Optical to Far-IR Mirror and Telescope TechnologyPI: H. Philip Stahl (MSFC)



LiteBIRD-specification lowfrequency detector array mounted in a dilution refrigerator for test

Significance: May enable future Cosmic Microwave Background (CMB) missions, e.g. LiteBIRD Project Title: Technology Development for LiteBIRD and other CMB Missions PI: Adrian T. Lee (UC Berkeley)



Evolution of microshutter arrays: A 128×64 FORTIS pilot on the left, a 365×171 JWST array in the center, and a space-qualified 736×384 next-gen array on the right Significance: May enable sparse-field multi-object spectroscopy for future strategic and other missions Project Title: Scalable Microshutter Systems for UV, Visible, and IR Spectroscopy PI: Paul Scowen (GSFC)



100-mm cross-strip anode fabricated with high-temperature co-fired ceramic (HTCC, 800°C) for applications in open-face and sealed-tube Micro-Channel-Plate MCP) detectors with high spatial resolution (20 μ m) and high event rates (> 5MHz) to address some HWO technology gaps. **Significance:** May enable UV/Visible light detection for future strategic missions such as an IR/O/UV Great Observatory

Project Title: High-Performance Sealed-Tube Cross-Strip (XS) Photon-Counting Sensors for UV-Vis Astrophysics Instruments

PI: Oswald Siegmund (UC Berkeley)



Completed 200-mm diameter, back-illuminated CCD wafer

Significance: Future strategic X-ray observatories require soft X-ray (sub-keV) spectral response close to the Fano limit over large detector areas and for multiple detectors Project Title: Optimized Soft X-ray Sensors for Strategic X-ray Astrophysics Missions: Achieving TRL 5 PI: Christopher Leitz (MIT/LL)



64-pixel prototype broadband far-IR Microwave Kinetic Inductance Detector (MKID) with multiplexed readout electronics

Significance: New broadband, scalable, and generalizable far-IR detector technology with compact and efficient data acquisition applicable to future NASA missions Project Title: Far-IR Detector Solutions for Low Noise, Large Format, Direct Absorption Kinetic Inductance Detector Array PI: Jason Austermann (NIST)



Test setup for LmAPD detectors

Significance: Ultra-low-noise detectors may enable spectroscopy of extrasolar planets Project Title: Photon-Counting Near-IR Linear-mode Avalanche-Photo-Diode (LmAPD) Arrays for Ultra-low Background Space Observations PI: Michael Bottom (U. of Hawaii) Page 17 of 25



Timepix4 512 x 448 pixelated readout ASIC on a 300-mm wafer

Significance: Four-side-buttable lowpower readout chips may enable future far-UV missions with large focal planes Project Title: Large-Format, High-Dynamic-Range UV detector using MCPs and Timepix4 readouts PI: John Vallerga (UC Berkeley)



Delta-doped CMOS image sensors

Significance: Astro2020 science goals require multi-gigapixel mosaic focal planes with largeformat CMOS detectors (8k×8k), low-noise (< 2.5 e-), small pixels (5-10 μm), broadband UV/Optical/IR response (>50% Quantum Efficiency, QE), and visible-blind near-UV detectors with high QE for 200-400 nm **Project Title:** High Performance Far-UV. Near-UV, and UV/Optical CMOS Imagers

Project Title: High Performance Far-UV, Near-UV, and UV/Optical CMOS Imagers **PI:** Michael Hoenk (JPL)



SQUID-based multiplexer arrays with Two-Level Switching (TLS) optimized for bolometer readout Significance: Advancing time-domain multiplexing (TDM) readout for large-format Transition-Edge-Sensing (TES) bolometers could enable or enhance the next far-IR Great Observatory Project Title: Advancing Readout of Large-Format Far-IR Transition-Edge Sensor Arrays PI: Karwan Rostem (GSFC)



Coefficient of Moisture Expansion (CME)/Creep specimens for testing in ultra-stable test bed Significance: Ultra-stability and precision (~10 pm) may enable the next IR/optical/UV Great Observatory Project Title: Ultra-Stable Structures: Development and Characterization Using Spatial Dynamic Metrology PI: Babak Saif (GSFC)



Short-wavelength (25 μ m) absorber/inductor design

Significance: Large-format arrays of sensitive far-IR detectors will enable space-based spectroscopy many orders of magnitude more sensitive than previous facilities **Project Title:** Ultrasensitive Far-IR Kinetic Inductance Detector (KID) Arrays for Space **PI:** Steven Hailey-Dunsheath (California Institute of Technology)



A very-low-blaze angle grating prototype made for the FORTIS sounding rocket (PI: McCandliss) Significance: Very-low-blaze angle (< ~1 deg) UV gratings enable spectroscopy for missions such as FORTIS, as well as Explorers, Probes, and Flagships like the Habitable Worlds Observatory **Project Title:** UV Spectroscopy for the Next Decade Enabled Through Nanofabrication Techniques **PI:** Randall McEntaffer (PSU)



Flight-like JPL package for hybridized 12x84-pixel kinetic inductance detector (KID) array bonded to a matching GSFC microlens array (on the back side of the KID) with a zoom-in showing individual pixels (right)

Significance: Extremely sensitive far-IR detectors may enable future missions Project Title: Ultrasensitive Far-IR Kinetic Inductance Detector (KID) Arrays: Maturation for Flight PI: C. Matt Bradford (JPL) Page 24 of 25



Atomic-Layer-Deposition (ALD) encapsulation of Physical-Vapor-Deposition (PVD) mirror coatings for improved stability, same as used for the SPRITE and Aspera missions Significance: Advanced coatings may enable future far-UV missions Project Title: High-Performance, Stable, and Scalable UV Aluminum Mirror Coatings Using ALD PI: John Hennessy (JPL)